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# Low noise cryogenic system for the measurement of the Casimir energy in rigid cavities

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## Abstract

We report on preliminary results on the measurement of variations of the Casimir energy in rigid cavities through its influence on the superconducting transition of in-cavity aluminium (Al) thin films. After a description of the experimental apparatus, we report on a measurement in the presence of thermal photons, discussing its implications for the zero-point photons case. Finally, we show preliminary results for the zero-point case.

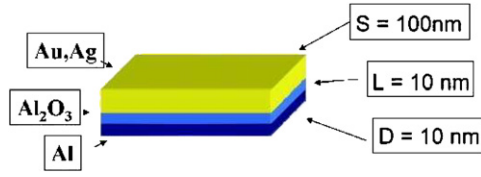
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(Some figures in this article are in colour only in the electronic version)

## 1. Introduction

In spite of great theoretical and experimental advances [1, 2], the Casimir effect still faces important unsolved questions. In gravitation and cosmology, the problem of reconciling the vacuum energy density and its interaction with the gravitational field, known as the cosmological constant problem [3, 4], remains unsolved. In flat spacetime, theoretical disagreements still remain on the dependence of the Casimir energy (and force) on the geometrical shape of the bodies and on the possibility that some geometries exhibit positive energy and repulsive force [1]. All experiments performed so far have been addressed to measuring the forces between (two) separate bodies, not directly the (variation of) energy in a rigid body. Apart from the fundamental interest of a direct measurement of the energy, the use of rigid cavities could make it easier to test different shapes. The interest is not only theoretical:

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**Figure 1.** Schematic of a Al–Al<sub>2</sub>O<sub>3</sub>–Au cavity.

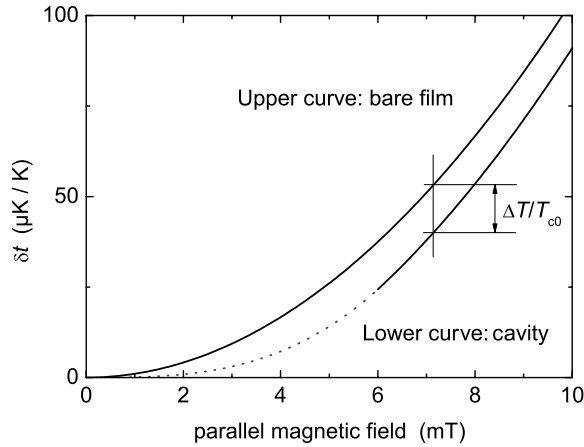
repulsive forces might build a bridge between the Casimir effect and nanotechnology. On this basis, we recently proposed the experiment Aladin2, to measure the variation of the Casimir energy in a rigid cavity, occurring when a thin metallic film, constituting one of the plates of the cavity (the other plate being made of a normal metal), becomes superconducting [5]. The experiment is made possible by the observation that, under suitable conditions, the variation of vacuum energy in the transition (due to the change in the film reflectivity)  $\Delta F^{\text{cav}}$  can be comparable with the condensation energy  $E_{\text{cond}}$  of the superconducting film, and therefore may sensibly affect the transition itself [5, 6]. In particular, the parallel critical field  $H_{\parallel}(T)$  required to destroy the superconductivity of the film can be remarkably different from the critical field of a bare film so that, to prove this effect, the experiment measures the difference in the values of  $H_{\parallel}(T)$  for a film which is the part of a Casimir cavity with respect to an equal but bare film. The magnitude of the effect depends also on the roughness of the plates and on their conductivity, and hence, in principle, our method might be used to probe the dependence of the Casimir effect on these factors, which are currently being investigated experimentally. We remark that our approach, being based on rigid cavities, could be easily extended to measure the energy of different shapes, like for example granular superconductors. Finally, we note that the scheme of our experiment arose as a solution to the problem of modulating the Casimir energy in a rigid cavity [7], while introducing in the system the minimal amount of energy. This goal is the part of a more speculative project, aimed at verifying whether or not the Casimir energy gravitates, by weighing a Casimir apparatus whose energy is modulated in time [7].

## 2. Expected signal

The Casimir cavity, shown in figure 1, is composed of a thin Al superconducting layer of thickness  $D$ , an intermediate dielectric thin layer (Al<sub>2</sub>O<sub>3</sub>) of thickness  $L$  and a third thick metallic layer made of gold (Au) or silver (Ag). (See below for more details.) The expected effect is shown in figure 2 where the shift  $\delta t = 1 - t$  of the reduced critical temperature  $t = T/T_{c0}$  is plotted against the critical field  $H_{\parallel}$ ;  $T_{c0}$  is the zero-field transition temperature. The upper curve is for a bare superconducting thin film, while the lower curve is for the Casimir cavity. For a bare film with  $1 - t \ll 1$ , as is our case, the critical field  $H_{\parallel}$  follows the well-known equation [8],

$$H_{\parallel} = H_0 \sqrt{24} \frac{\lambda(0)}{D} \sqrt{1 - t}, \quad (1)$$

which neglects the effect of nucleation, which applies to a very thin film. Here,  $\lambda$  is the penetration depth and  $H_0$  is the bulk zero-temperature critical field. On inverting equation (1), one obtains a parabolic dependence of  $\delta t$  on  $H_{\parallel}$ , corresponding to the upper curve of figure 2. In a cavity, the dependence is expected to differ from the bare-film case in the region  $\delta t \ll 1$ , corresponding to low applied fields. As shown in figure 2, the expected difference  $\Delta T$  in the



**Figure 2.** Simulation of the expected signal for a bare thin Al film of thickness  $D = 14$  nm (upper curve) and for a cavity composed of a similar Al film, covered by a 6 nm dielectric layer and a 100 nm Au mirror (lower curve).

transition temperature for low applied fields is in the range of a few  $\mu\text{K}$  (for higher values of the field, not shown in the plot, the two curves approach each other) [5, 6, 9]. In the region of very small applied fields, a deviation from the parabolic curve is expected by continuity arguments (dashed line in figure 2); in this latter region, the change in free energy becomes comparable with the condensation energy and no perturbative calculation can be done. The choice of the material is crucial, and a compromise is usually required to encompass various needs. On the one hand, low  $T_c$  materials (low condensation energy) are preferable because their transition is more perturbed by the vacuum energy effects [6]. On the other hand, the longer cooling times typically required when working with (very) low  $T_c$  materials could make it difficult to increase the statistics. In this first experiment, Al has been chosen as a good compromise. Another important parameter is the plasma frequency  $\Omega_p$  of the normal plate; larger values of  $\Omega_p$  imply a larger conductivity and ensure a better confinement of the photons within the cavity. It was shown in [6] that a very good material could be Be ( $\Omega_p^{\text{Be}} = 18.9$  eV), but difficulties in obtaining cavities with this material led us to use Au ( $\Omega_p^{\text{Au}} = 9.0$  eV) or Ag ( $\Omega_p^{\text{Ag}} = 8.2$  eV), resulting in a loss of approximately a factor of 2 in the expected signal [6]. The values of other electrical parameters, such as the electrons mean free paths both in the normal metal and in the superconductor, do not affect significantly the expected signal [6]. While the signal does not depend on the area of the plates, the optimal thickness  $D$  of the superconducting metal turns out to be a few nanometres. Numerical calculations show that the variation of vacuum energy in the transition has the following dependence on the gap  $L$ :

$$\Delta F_E^{(C)}(L) \propto \frac{1}{1 + (L/L_0)^\alpha}, \quad (2)$$

where  $\alpha = 1.15$  and  $L_0 \simeq 10$  nm. As we see, as soon as  $L$  becomes smaller than  $L_0$ ,  $\Delta F_E^{(C)}(L)$  approaches a finite value and therefore, once  $L$  is in the region of a few nanometres, no substantial gain can be obtained by using smaller gaps. It can be shown that the signal is practically independent of the dielectric material, and we chose the Al native oxide  $\text{Al}_2\text{O}_3$ . In order to overcome spurious effects due to different alignment of the cavity and the bare film with respect to  $H$ , the cavity and the bare film are placed on the same chip.

### 3. Sample preparation, cryogenic apparatus and measurement method

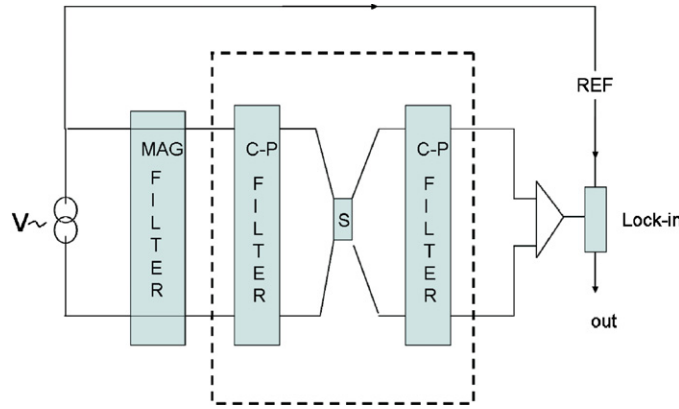
The films and cavities, made by IPHT (Institute for Photonic Technology, Jena, Germany), are obtained by depositing the Al on the whole wafer (silicon, 3 inch diameter). The only difference between the bare thin film and the cavity is the presence of a metallic layer above the dielectric layer for the cavity. Finally, the chip is cut in samples that include, within a distance of less than 1 mm, several films and cavities. Considering our parameters, as we reported elsewhere [9, 11], the applied magnetic field  $H$  must be very homogeneous, to within  $10^{-4}$  over the distance  $d$  between the cavity and the film ( $d \approx 0.5$  mm). This requirement has been fulfilled by a home-made coil system composed of a main coil and two lateral correction coils. The alignment of the field and sample is performed by mounting rigidly the sample holder and the coil system. If  $\theta$  is the angle between the magnetic field and the sample, following [8], for small angle  $\theta$ , a perpendicular component of the field adds a linear term in the dependence  $\delta t(H)$ , resulting approximately in

$$\delta t \approx \frac{D^2}{24\lambda^2(0)H_0^2} H^2 + \frac{\sin\theta}{H_0} H. \quad (3)$$

Since the film and the cavity are deposited at close distance on the same chip, and since the magnetic field is very homogeneous, possible differences in the values of  $\theta$  and  $H$  between the film and cavity can be neglected and therefore a non-vanishing  $\theta$  just yields a common additive term which does not affect the differential measurement.

We used films with different areas, ranging from about  $20 \times 20 \mu\text{m}^2$  to  $100 \times 100 \mu\text{m}^2$ , to verify that the effect does not depend on the area. Inspection of the samples with the atomic force microscope showed that the actual thicknesses of the Al layer and that of the oxide gap are of 14 nm and 6 nm, respectively. The roughness is about 1 nm for both. Measurements of different samples do not show any significant difference. The cryogenic system is based on the HelioxVL  $^3\text{He}$  commercial cryostat from Oxford Instruments, reaching a base temperature of 250 mK, inserted in a home-made dewar equipped with magnetic screening, to isolate the sample from undesired environmental fields. The usual ferrite filters at 300 K and 1 K-pot levels and a cascade of two isolation copper-powder filters [10] at the level of the 1 K-pot and the  $^3\text{He}$ -pot guarantee an efficient EM screening. The measurement method for the film and cavity resistances is a standard four-wire measurement, as shown in figure 3. The temperature is measured by a Cernox CX-1070-BR as a part of a Wheatstone bridge. Both measurements are done with lock-in technique. The actual measurement of  $R(T)$  is performed by fixing the external field, driving the temperature from low to high and measuring  $R(\tau)$  and  $T(\tau)$  ( $\tau$  is time) at the sample during the transition. The cool down is always performed in the zero magnetic field to avoid magnetic flux trapping. Each single transition measurement takes approximately 20 min.  $R(T)$  is finally recovered to evaluate the shift with respect to the zero-field case. The Al resistivity ratio of the cavities and films analysed in the present work is  $R(T = 300 \text{ K})/R_0(T = 1.6 \text{ K}) \simeq 2$ , compatible with the values quoted in the literature, corresponding to a mean free path of about 15 nm [12]. The samples' resistances vary from 200  $\Omega$  to 350  $\Omega$ . The transition temperature at zero field,  $T_{c0}$ , is taken at the maximum of the derivative of  $R(T, H = 0)$  and used as the normalization factor in  $\delta t$ . In order to minimize the error in the measurement,  $\delta t(H_{\parallel})$  is obtained as the difference of temperature not just at a given point on the transitions with and without the applied field (i.e. for a given resistance  $R$ ), but as the average of the differences for all points of the transition (i.e. for all values of  $R$ ):

$$\delta t = \frac{1}{T_{c0}} \cdot \overline{T(R, H_{\parallel} = 0) - T(R, H_{\parallel})}. \quad (4)$$



**Figure 3.** Schematic of the setup used for the measurement of sample resistance: the dashed box indicates the cryogenic environment, housing the sample and the copper-powder filters for the four measurement lines. The driving electronic and the mixer, which are at room temperature, are also shown.

The average is taken over the points with  $0.2 < \frac{R}{R_N} < 0.8$ , where  $R_N$  is the resistance in the normal state.

#### 4. Measurement in the presence of thermal photons

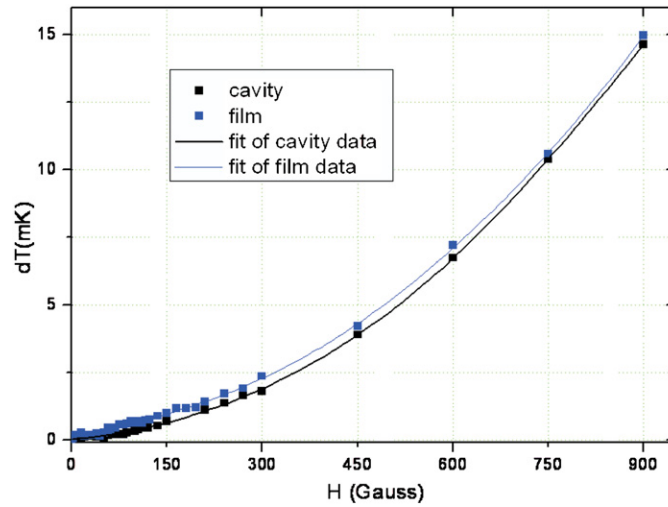
In order to test the cavity efficiency in removing photons affecting the transition, we performed a first measurement in the presence of room-temperature thermal photons. This has been done by removing the cascade of copper-powder filters and allowing the external rf and microwave photons enter the vacuum chamber. As pointed out in [6], the important region of frequencies  $\nu_{\text{eff}}$  contributing to our effect is around a few times  $2k_B T_c/h$ . In the absence of a detailed theory, as a first guess one can imagine that the presence of extra thermal photons from the environment at temperature  $T_E = 300$  K gives rise to a temperature shift  $(\Delta T)'$  that is larger than the previous  $\Delta T$  by a factor  $M$  equal to the ratio between the black-body densities  $E_{T_E}(\nu_{\text{eff}})$  and  $E_{T_c}(\nu_{\text{eff}})$  of electromagnetic energy at frequency  $\nu_{\text{eff}}$ , at temperatures  $T_E$  and  $T_c$ , respectively. Upon taking, to be definite,  $h\nu_{\text{eff}} = 10k_B T_c$ , we obtain

$$M = \frac{E_{T_E}(\nu_{\text{eff}})}{E_{T_c}(\nu_{\text{eff}})} \approx \frac{2}{e^{\frac{10T_c}{T_E}} - 1} \approx \frac{T_E}{5T_c} \approx 40. \quad (5)$$

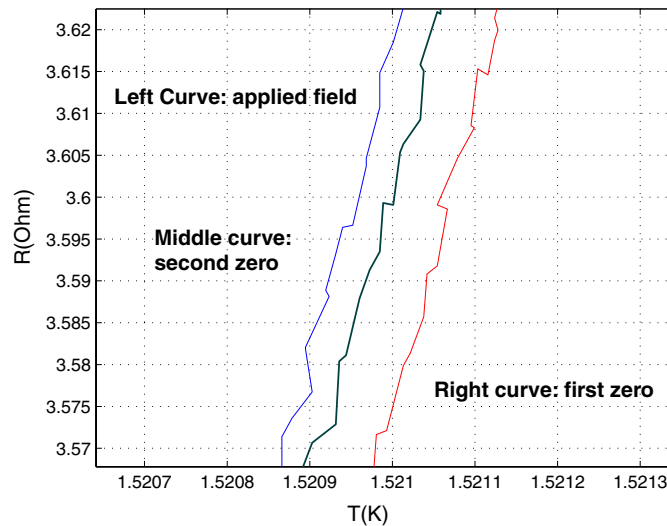
From this estimate, we expect that this measurement (which, apart from the filters, is the same as the one described earlier) should produce a signal similar to the one described in figure 2, with  $\Delta T$  increased by a factor of some tens. As shown in figure 4, this is indeed what we found in our measurements: the curves relative to the bare film and the cavity, for low  $H$ , are separated by a gap which reaches a maximum of about  $(\Delta T)' = 300 \mu\text{K}$ . We point out that, although it must be considered as a rough estimate, this measurement suggests a value of  $\Delta T$  in the final experiment close to the most conservative expectation of  $\Delta T = 10 \mu\text{K}$  (or even less).

#### 5. First results in the measurement of Casimir energy variations

In the presence of copper-powder filters, the major source of noise is the electronic noise at the read-out amplifier. This noise has a ‘fast’ component (with the time scale of one

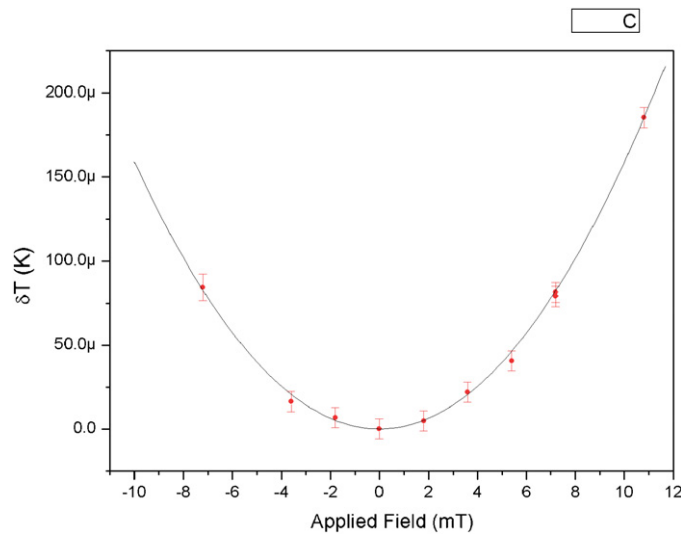


**Figure 4.** Results for the measurements without copper-powder filters (the lower curve is for the cavity, the upper for the bare film). The behaviour is qualitatively similar to the theoretical prediction in figure 2, but the shift is enhanced from the presence of  $T_E = 300$  K thermal photons: the maximum shift in temperature is about  $300 \mu\text{K}$ , much higher than  $1\text{--}10 \mu\text{K}$  expected for the zero-point case.



**Figure 5.** Example of a measurement triplet: a zoom of the resistance versus temperature for two transitions in zero field and one with  $\mu_0 H = 7.2$  mT. The mean shift is  $80 \mu\text{K}$ .

second) that determines a statistical error of a few  $\mu\text{K}$ , and a slow thermal drift (linear in time, about  $-50 \mu\text{K}$  per hour) that produces a  $\delta t$  proportional to the time elapsed between two measurements. To correct for the thermal drift, at equal time intervals before and after each measurement with the applied field, we perform two measurements in zero field, and then we take for  $\delta t$  the average of the two  $\delta t$ . We performed the measurement on three different samples, without finding significant differences. An example of such a measurement is shown



**Figure 6.** Experimental low-field points for a bare film. The solid line is a parabola obtained from a fit of high-field measurements (data not shown). The error bars are of  $6 \mu\text{K}$ , as estimated in the text. The good agreement indicates that there are no unexpected sources of noise.

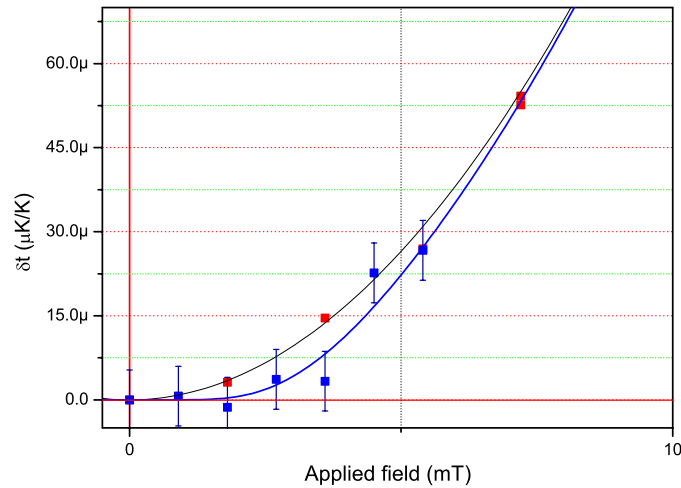
in figure 5: the intermediate curve and the right curve were taken in zero field, while the left curve is with an applied field of 7.2 mT. It should be noted that the shifts in temperature for the three curves are only a few tens of  $\mu\text{K}$ . This procedure is repeated for all applied fields.

The data for a typical bare film are reported in figure 6. The parallel field is applied in opposite directions to rule out possible hysteresis. We find that the theoretical parabolic dependence of  $\delta t$  on the applied field is recovered for fields as low as a few mT, corresponding to changes of the critical temperature of about  $10 \mu\text{K}$ . The sensitivity has been roughly estimated to be about  $6 \mu\text{K}$  by comparing several measurements of  $\delta t$  under the same conditions, performed at some days of distances. It is important to stress that the solid line was obtained by fitting high-field measurements (data not shown), which are less affected by the possible presence of EM noise or field–film misalignment. The good agreement of the solid line with the low-field measurements indicates that there are no unexpected sources of noise in the low-field region.

An example of a complete measurement, for one of our samples, is reported in figure 7. Other samples do not show significant differences at the present level of sensitivity. The upper curve is the same bare-film parabola of figure 6, while the lower curve is a fit on the cavity data. To help reading the figure, the bare-film data are reported without errors. The cavity curve lies below the bare-film curve, and shows a non-parabolic behaviour, in qualitative agreement with theoretical expectations. Moreover, the maximum displacement between the two curves, of about  $7 \mu\text{K}$ , is of the expected magnitude. However, it is also comparable with the present sensitivity and therefore the result of these measurements, while encouraging, is far from being conclusive. In addition to this, the two curves appear to merge much faster than expected theoretically, and the reason for this is presently not clear.

In order to increase the sensitivity, a new cryogenic apparatus is being built, allowing to reach lower temperatures (about 20 mK) and to use softer superconductors. We expect





**Figure 7.** Results of the final measurements. The lower curve is the cavity, the upper is for the bare film. (Error bars are  $6 \mu\text{K}$ , see text.) The cavity shows a different behaviour at very low fields (as expected from the theory).

that the signal-to-noise ratio will improve, as discussed in section 2, to a sufficient extent to definitively confirm or disprove the foreseen effect.

## References

- [1] Bordag M, Mohideen U and Mostepanenko V M 2001 *Phys. Rep.* **353** 1
- [2] For most recent experiments:  
Castillo-Garza R, Chang C C, Jimenez D, Klimchitskaya G L, Mostepanenko V M and Mohideen U 2007 *Phys. Rev. A* **75** 062114  
Bressi G, Carugno G, Onofrio R and Ruoso G 2002 *Phys. Rev. Lett.* **88** 041804  
Decca R S, López D, Fischbach E, Klimchitskaya G L and Krause D E 2007 *Phys. Rev. D* **75** 077101  
Iannuzzi D, Lisanti M and Capasso F 2004 *Proc. Natl Acad. Sci. USA* **101** 4019
- [3] Weinberg S 1989 *Rev. Mod. Phys.* **61** 1
- [4] Ishak M 2007 *Found. Phys.* **37** 1470
- [5] Bimonte G, Calloni E, Esposito G, Milano L and Rosa L 2005 *Phys. Rev. Lett.* **94** 180402
- [6] Bimonte G, Calloni E, Esposito G and Rosa L 2005 *Nucl. Phys. B* **726** 441
- [7] Calloni E, Di Fiore L, Esposito G, Milano L and Rosa L 2002 *Phys. Lett. A* **297** 328
- [8] Tinkam M 1975 *Introduction to Superconductivity* (New York: McGraw-Hill)
- [9] Bimonte G, Born D, Calloni E, Esposito G, Hübner U, Il'ichev E, Rosa L, Tafuri F, Scaldaferrì O and Vaglio R 2007 *Nucl. Instrum. Methods A* **572** 515
- [10] Martinis J M, Devoret M H and Clarke J 1987 *Phys. Rev. B* **35** 4682
- [11] Bimonte G, Born D, Calloni E, Esposito G, Hübner U, Il'ichev E, Rosa L, Tafuri F, Scaldaferrì O and Vaglio R 2006 *J. Phys. A: Math. Gen.* **39** 6153
- [12] Van Son P C, Romijn J, Klapwijk T M and Mooij J E 1984 *Phys. Rev. B* **29** 1503